

Fast and accurate field reconstruction of Thermal-Large Eddy Simulation (T-LES) by Deep Learning

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Context: Issues

- Only one side of the solar receiver is heated
- Implies high temperature gradient

Flows in the receiver

- The flows in the new generation solar receivers are turbulent and anisothermal
- Thermal exchanges at the **boundary** define the **energy efficiency**

Scientific bottleneck

• Understand the coupling between **velocity** and **temperature** in highly anisothermal flows

Context: Simulations

DNS

T-LES

- Modeling of strongly anisothermal turbulence using Thermal Large Eddy Simulation (T-LES)
- T-LES only simulates the large scales, and models the small scales

Context: Simulations

- Development of deconvolution techniques to accurately reconstruct fields
	- ⇒ Estimation of RMS temperature for performance of T-LES models against DNS
	- ⇒ Estimation of RMS heat flux for estimation of thermomechanical constraints

Context: Study case

- Hot wall: concentrated sunlight, cold wall: insulated
- Fluid: air
- Regular mesh in the x and z directions
- Hyperbolic tangent mesh in the direction normal to the walls y
- Half height of the channel $h = 15$ mm

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General principles

- Training algorithms to recognize patterns in data and make predictions or decisions based on that data.
- Takes the form of a non-linear optimization problem over labeled data *(i.e.* for each input, there's an expected output)
- Stacking layers with non linearities in between them
- Many applications: Image generation, fraud detection, text translation, artistic tools, etc... .

Convolutional neural network

Convolutional Neural Networks (CNNs)

- Type of architecture with a learnable (optimizable) convolutional kernel
- Kernel can be any dimention, we can learn

Convolutional neural network layer

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Super-resolution

- The LES filter is unknown
- Learning of an inversion operator for filtering on the temperature field:

$$
T_{DNS} \approx G^{-1} * T_{LES}
$$

• The neural network learns the correction on the LES field

 $T_{\text{pred}} = T_{\text{LES}} + f_{\text{CNN}}(T_{\text{LES}}) \approx T_{\text{DNS}}$

- Anisothermal DNS, mean friction Reynolds number $Re_{\tau} = 180$, and Prandtl $Pr = 0.76$ after statistical convergence
- 17 DNS snapshots spaced by $\Delta_t^+ = 7.76 \times 10^{-3}$
- Interpolate from a fine mesh of $(384, 384, 266)$ points to $(48, 48, 52)$ points
- Filter using a weighted top-hat filter
- Filtered DNS snapshots then serve us as input to our network. $14/31$

Benchmark

• We compare the performance of our motel to an already existing method developped by Van Cittert The inverse of a convolution filter G assumed invertible, writes

$$
G^{-1} = (\mathcal{I} - (\mathcal{I} - G))^{-1}, \tag{1}
$$

$$
= \lim_{p \to \infty} \sum_{i=0}^{p} (Z - G)^{i}.
$$
 (2)

We take a $p = 6$ approximation to this converging Neumann series [\(Stolz and](#page-30-1) [Adams \[1999\]](#page-30-1) recommend $p = 5$).

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- 17 snapshots are split into 13 for training, 4 for validation and 1 for test
- We learn over patches, and increase procedurally the sampling space
- At each fixed height, with patches of size 16, we have the following number of examples

$$
\frac{n_x}{16} \times \frac{n_z}{16} \times \text{flip}_x \times \text{flip}_z \times \text{rotation}_{xz} \times n_{\text{train}} = 936
$$

- Previous technique domain increase procedure
- Enables learning harder distributions of flow
- $Re_{\tau}_{hot} = 105$, $Re_{\tau}_{cold} = 260$

- Neural network manages to quickly learn and generalize
- Jumps are due to increase of the learning domain

Results: Validation slices, T_{valido}

Spacial gradients in our method are **slightly high** 20/31

Results: Validation slices, T_{valid_1}

Results: Validation slices, T_{valid}

Results: Validation slices, T_{valid}

• Slight overshoot of the RMS on the hot side $24/31$

 285

 345

 $T(K)$

465

 525

586

 -405

• Good prediction of the temperature probability distribution 27/31

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In conclusion

- Our method learns and generalizes the small scales reconstruction even for snapshots taken farther in time for RMS, densities and slices
- Overshot spacial gradients can be mitigated through an additional filtering layer
- Good tool for the study of flows inside high temperature solar receivers

Future work

- Learning over different meshes
- Learning over **distribution** of data
- Generalizing over different Reynolds number flows

Thank you for your attention

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C. J. Lapeyre, A. Misdariis, N. Cazard, D. Veynante, and T. Poinsot. Training convolutional neural networks to estimate turbulent sub-grid scale reaction rates. Combustion and Flame, 203:255–264, May 2019. doi: 10.1016/j.combustflame.2019.02.019. URL <https://hal.archives-ouvertes.fr/hal-02072920>.

S. Stolz and N. Adams. An approximate deconvolution procedure for large-eddy simulation. Phys. Fluids, 11(7):1699–1701, 1999.